

The role of light polarization and related effects in Maxwell fish eye refractive profile

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PhD school and workshop “ASPECTS OF SYMMETRY”
(November 8-12, 2021)

Based on the following papers:

1. M. D., Z. Gevorkian, & A. Nersessian (2021). Maxwell fish eye for polarized light. [Physical Review A, 104\(5\), 053502.](#)
2. Z. Gevorkian, M.D., & A. Nersessian (2020). Extended symmetries in geometrical optics. [Physical Review A, 101\(2\), 023840.](#)

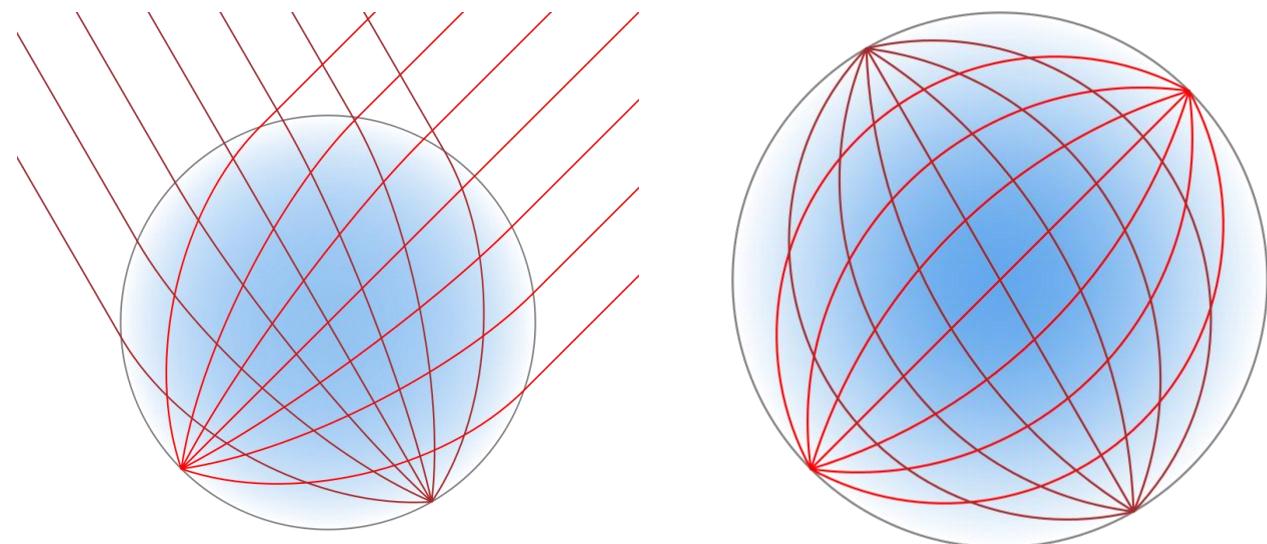
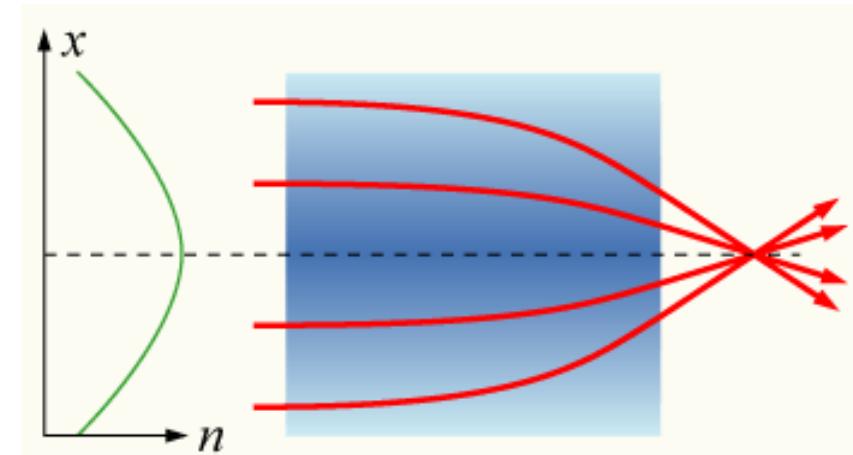
Structure

- Refractive index profiles, the “Maxwell fish eye”
- Hamiltonian formalism
- Construction scheme of optical Hamiltonians and refraction indices with similar symmetry algebras
- Coulomb-Fish eye “duality”
- Construction of MFE profile for polarized light
- Trajectories for polarized light
- Deformations of the trajectories

Refractive index profiles, the “Maxwell fish eye”.

- if the refractive index of a medium is not constant but has some sort of spatial dependency, the material is known as a gradient-index or GRIN medium
- the crystalline lens of the human eye is an example of a GRIN lens with a refractive index varying from about 1.406 in the inner core to approximately 1.386 at the less dense cortex
- Luneburg lens: $n(\mathbf{r}) = \sqrt{2 - (\mathbf{r}/R)^2}$
- Maxwell fish eye lens (n_0 and R are constants):

$$n(\mathbf{r}) = \frac{n_0}{1 + (r/R)^2}$$



Geometry

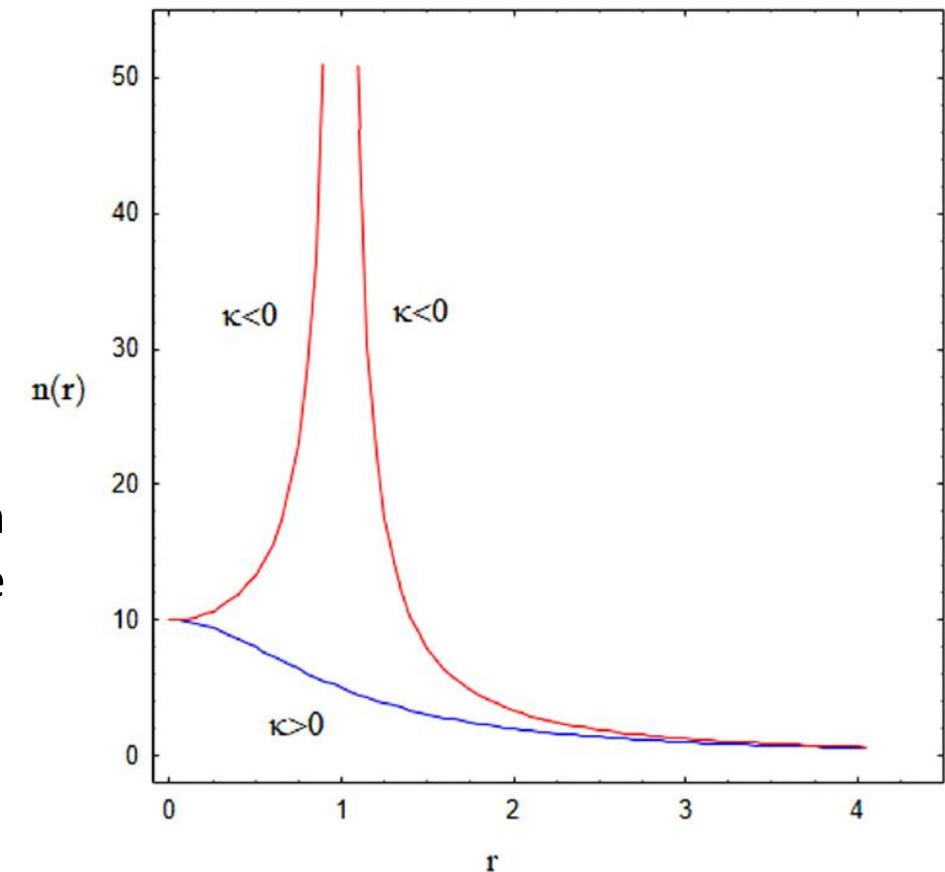
- action of a system on a three-dimensional curved space equipped with the “optical metrics” of Euclidean signature $d\tilde{l}^2 = n^2(\mathbf{r})d\mathbf{r} \cdot d\mathbf{r}$

$$S_{\text{Fermat}} = \frac{1}{\lambda_0} \int d\tilde{l}, \quad d\tilde{l} := n(\mathbf{r}) |d\mathbf{r}/d\tau| d\tau$$

- MFE is defined by the metrics of (three-dimensional) sphere or pseudosphere (under pseudosphere we mean the upper (or lower) sheet of the two-sheet hyperboloid).

$$n(\mathbf{r}) = \frac{n_0}{|1 + \kappa \mathbf{r}^2|}, \quad \kappa = \pm \frac{1}{4r_0^2}$$

- the symmetries of the system which describe the propagation of light in a particular medium are coming from the symmetries of the optical metrics of that particular medium
- $\text{SO}(3)$ and $\text{SO}(4)/\text{SO}(3.1)$



Hamiltonian formalism

$$\mathcal{H}_0 = \alpha(\mathbf{p}, \mathbf{r})\Phi = \alpha(\mathbf{p}, \mathbf{r}) \left(\frac{\mathbf{p}^2}{n^2(\mathbf{r})} - \lambda_0^{-2} \right) \approx 0 \quad \Phi := \frac{\mathbf{p}^2}{n^2(\mathbf{r})} - \lambda_0^{-2} = 0.$$

$$\{x_i, p_j\} = \delta_{ij}, \quad \{p_i, p_j\} = \{x_i, x_j\} = 0,$$

$$\frac{df(r, p)}{d\tau} = \{f, \mathcal{H}_0\} = \{f, \alpha\}\Phi + \alpha\{f, \Phi\} \approx \alpha\{f, \Phi\}. \quad \alpha = \frac{n^2(\mathbf{r})}{p + \lambda_0^{-1}n(\mathbf{r})} \quad \Rightarrow \quad \mathcal{H}_{\text{Opt}} = p - \lambda_0^{-1}n(r)$$



$$\frac{d\mathbf{p}}{dl} = \lambda_0^{-1} \nabla n(\mathbf{r}) \quad \frac{d\mathbf{r}}{dl} = \frac{\mathbf{p}}{p}$$

Construction of optical Hamiltonians with similar symmetry algebras

- given the Hamiltonian $H = \frac{p^2}{2g(r)} + V(r)$ after fixing the energy surface $H = E$, (according to Maupertuis principle) we can relate the trajectories with the optical Hamiltonian \mathcal{H}_0 and with the refraction index $n(r) = \sqrt{2g(r)(E - V(r))}$
- \mathcal{H}_0 , \mathcal{H}_{Opt} as well as $n(r)$ inherit all the symmetries and constants of motion of the Hamiltonian
- canonical transformations preserve the symmetries of the Hamiltonians and their level surfaces



we can to construct the physically non-equivalent optical Hamiltonians (and refraction indices) with the identical symmetry algebra

Coulomb-Fish eye “duality”: Coulomb problem

$$H_{\text{Coul}} - E := \frac{p^2}{2} - \frac{\gamma}{r} - E = 0, \quad \Rightarrow \quad n_{\text{Coul}} = \lambda_0 \sqrt{2(E + \gamma/r)}, \quad \text{where } \gamma > 0.$$

- **Integrals of motion:** $L = r \times p, \quad A = L \times p + \gamma \frac{r}{r}$
- **Algebra formed:** $\{A_i, A_j\} = -2\epsilon_{ijk}H_{\text{Coul}}L_k, \quad \{A_i, L_j\} = \epsilon_{ijk}A_k, \quad \{L_i, L_j\} = \epsilon_{ijk}L_k$

Coulomb-Fish eye “duality”: Maxwell Fish eye

$$(\mathbf{p}, \mathbf{r}) \rightarrow (-\mathbf{r}, \mathbf{p})$$

$$r^2 - \frac{2\gamma}{p} - 2E = 0 \Rightarrow p - \frac{2\gamma}{r^2 - 2E} = 0 \Rightarrow n_{Mfe} = \frac{n_0}{|1 + \kappa r^2|}, \text{ where } \kappa := -\frac{1}{2E}, \quad \frac{n_0}{\lambda_0} := 2\epsilon\kappa$$

- **Integrals of motion:** $\mathbf{L} \rightarrow \mathbf{L}, \quad \mathbf{A} \rightarrow \frac{\mathbf{T}}{2\kappa}, \quad \mathbf{T} = (1 - \kappa r^2)p + 2\kappa(rp)\mathbf{r}$
- **Algebra formed:** $\{L_i, L_j\} = \epsilon_{ijk}L_k, \quad \{T_i, L_j\} = \epsilon_{ijk}T_k, \quad \{T_i, T_j\} = 4\kappa\epsilon_{ijk}L_k$

Inclusion of polarization

Inclusion of polarization in terms of Hamiltonian formalism means to preserve the form of the Hamiltonian and replace the canonical Poisson brackets by the twisted ones

$$\{x_i, p_j\} = \delta_{ij}, \quad \{x_i, x_j\} = s\varepsilon_{ijk}F_k(p), \quad \{p_i, p_j\} = 0,$$



F_k are the components of the Berry monopole \mathbf{F} :

$$\mathbf{F} := \frac{\partial}{\partial \mathbf{p}} \times \mathbf{A}(\mathbf{p}) = \frac{\mathbf{p}}{\mathbf{p}^3}$$

where

\mathbf{A} is the vector-potential of the “Berry monopole” i.e. the potential of the magnetic (Dirac) monopole located at the origin of momentum space

Rotation generators in this space :

$$\mathbf{J} = \mathbf{r} \times \mathbf{p} + s \frac{\mathbf{p}}{p}$$

Equations of motion:

$$\frac{d\mathbf{p}}{dt} = \lambda_0^{-1} \nabla n(\mathbf{r}), \quad \frac{d\mathbf{r}}{dt} = \frac{\mathbf{p}}{p} - \frac{s}{\lambda_0} \mathbf{F} \times \nabla n(\mathbf{r})$$

However, the above procedure, i.e. twisting the Poisson bracket with preservation of the Hamiltonian, violates the non-kinematical (hidden) symmetry of the system.

Construction of MFE profile for polarized light: MICZ-Kepler problem

- Coulomb problem in the presence of Dirac monopole (MICZ-Kepler problem)

$$H_{MICZ} = \frac{p^2}{2} + \frac{s^2}{2r^2} - \frac{\gamma}{r} \quad \{x_i, p_j\} = \delta_{ij}, \quad \{p_i, p_j\} = s\epsilon_{ijk} \frac{x_k}{r^3}, \quad \{x_i, x_j\} = 0$$

twisted Poisson brackets

- **Integrals of motion:** $\mathbf{J} = \mathbf{r} \times \mathbf{p} + s \frac{\mathbf{r}}{r}$ $\mathbf{A}_s = \mathbf{J} \times \mathbf{p} + \gamma \frac{\mathbf{r}}{r}$

- **Algebra:** same as in previous case with the replacement $(\mathbf{L}, \mathbf{A}) \rightarrow (\mathbf{J}, \mathbf{A}_s)$

Construction of MFE profile for polarized light: Deformed MFE

$$(\mathbf{p}, \mathbf{r}) \rightarrow (-\mathbf{r}, \mathbf{p})$$

$$H_{MICZ} = E \iff r^2 + \frac{s^2}{p^2} - \frac{2\gamma}{p} - 2E = 0 \Rightarrow$$

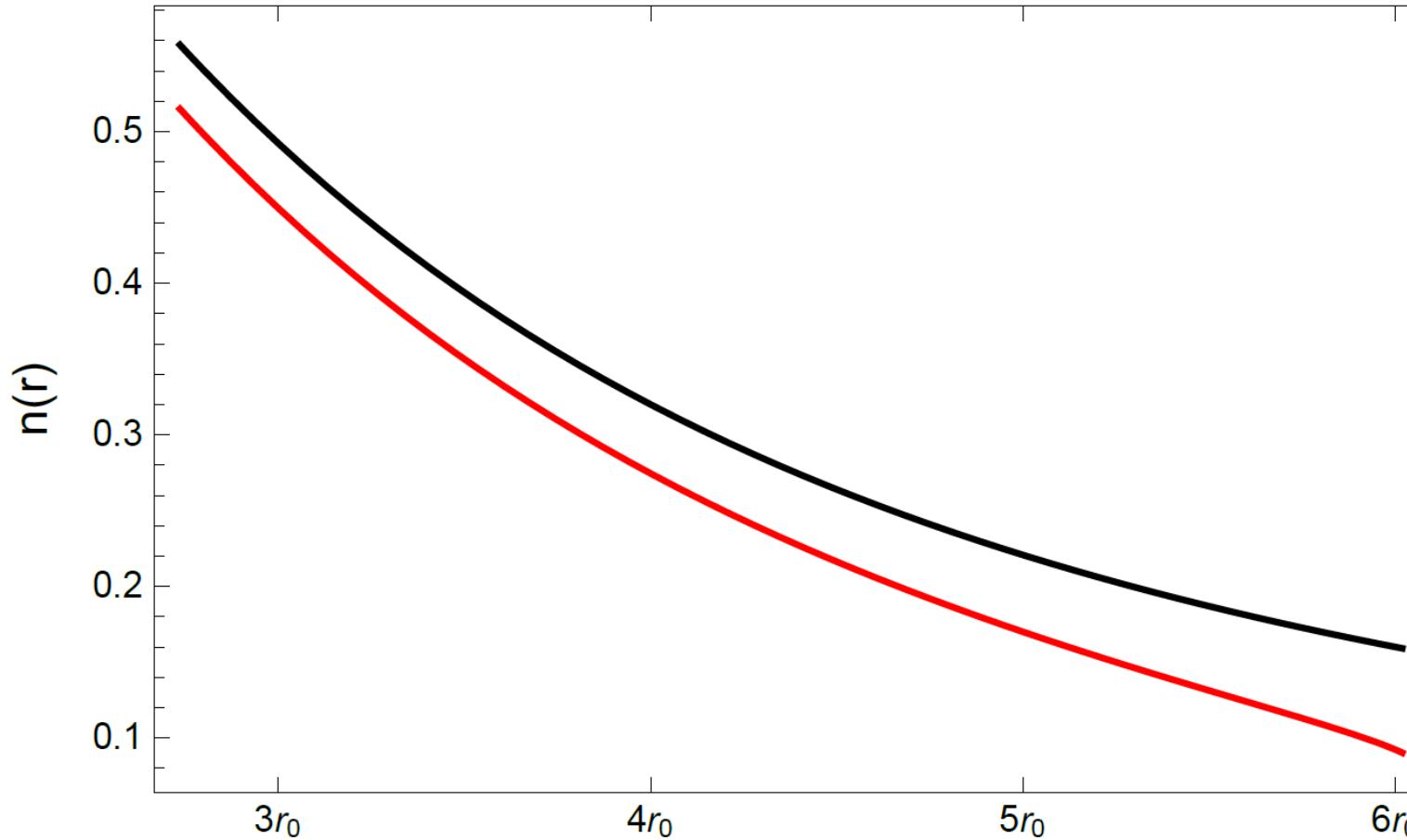
$$n_{Mfe}^s(\mathbf{r}) = \frac{n_{Mfe}(\mathbf{r})}{2} \left(1 + \sqrt{1 - \frac{4\kappa s^2 \lambda_0^2}{n_0} \frac{1}{n_{Mfe}(\mathbf{r})}} \right),$$

where $\kappa := -\frac{1}{2E}$, $\frac{n_0}{\lambda_0} := 2\epsilon\kappa$, $\epsilon = -\text{sgn}(r^2 + 1/\kappa)$

Construction of MFE profile for polarized light: Integrals of motion

- **Integrals of motion:** $\mathbf{J} \rightarrow \mathbf{r} \times \mathbf{p} + s \frac{\mathbf{p}}{p}$, $\mathbf{A}_s \rightarrow \frac{\mathbf{T}_s}{\kappa}$, $\mathbf{T}_s = \left(2 - \frac{n_0}{n_{Mfe}^s(\mathbf{r})}\right) \mathbf{p} + 2\kappa(rp)\mathbf{r} + \frac{2\kappa s}{n_{Mfe}^s(\mathbf{r})} \mathbf{J}$
- **Algebra:** same as the original MFE profile with the replacement $(\mathbf{L}, \mathbf{T}) \rightarrow (\mathbf{J}, \mathbf{T}_s)$

MFE profile for polarized light



$$n_{\text{Mfe}}^s(\mathbf{r}) = \frac{n_{\text{Mfe}}(\mathbf{r})}{2} \left(1 + \sqrt{1 - \frac{4\kappa\lambda_0^2 s^2}{n_0} \frac{1}{n_{\text{Mfe}}(\mathbf{r})}} \right)$$

\mathbf{r}

$\text{---} \quad s = 1$
 $\text{---} \quad s = 0$

- for $\kappa > 0$ we get a restriction of rays in the finite domain: $r \leq \sqrt{\frac{n_0^2}{4s^2\lambda_0^2\kappa^2} - \frac{1}{\kappa}}$
- note that spin appears along with the factor $\kappa\lambda_0^2 = (\lambda_0/2r_0)^2$
- in order to stay within the bounds of geometrical optics approximation, this factor must be reasonably small.
- therefore, the influence of the spin will be far more notable within certain range of distance from the core of the fish eye.
- the latter happens when the condition $4\kappa s^2 \lambda_0^2 / n_0 \approx n_{\text{Mfe}}(\mathbf{r})$ holds
- at these distances the refraction index in the presence of spin can be much smaller compared to the refraction index with zero spin

Trajectories

$$\mathbf{r} \cdot \mathbf{J} = s \frac{\mathbf{r}\mathbf{p}}{p}, \quad \mathbf{r} \cdot \mathbf{T}_s = \frac{n_0}{\lambda_0} \frac{\mathbf{r}\mathbf{p}}{p} \quad \Rightarrow \quad \mathbf{r} \cdot \left(\mathbf{J} - \frac{s\lambda_0}{n_0} \mathbf{T}_s \right) = 0 \quad \Rightarrow \quad \boxed{\text{ray trajectories are orthogonal to the axis} \\ \mathbf{E}_3 = \mathbf{J} - \frac{s\lambda_0}{n_0} \mathbf{T}_s}$$

⇓

the trajectories belong to the plane spanned by the following vectors:

$$\mathbf{E}_1 = \mathbf{T}_s \times \mathbf{J}, \quad \mathbf{E}_2 = \mathbf{E}_3 \times \mathbf{E}_1 = \left(\mathbf{J}^2 - s^2 \right) \left(\mathbf{T}_s - \frac{4s\lambda_0 \kappa}{n_0} \mathbf{J} \right) \quad : \quad \mathbf{E}_3 \cdot \mathbf{E}_2 = \mathbf{E}_3 \cdot \mathbf{E}_1 = 0$$

Then, from the expression $\mathbf{J} \cdot (\mathbf{r} \times \mathbf{T}_s)$ we immediately obtain the solution for the ray trajectories:

$$\mathbf{r} \cdot (\mathbf{T}_s \times \mathbf{J}) = \left(\mathbf{J}^2 - s^2 \right) \left(2 - \frac{n_0}{n_{\text{mfe}}^s} \right)$$

Trajectories: Polar coordinates

We introduce the following orthogonal frame

$$\mathbf{e}_i = \frac{\mathbf{E}_i}{|\mathbf{E}_i|} : \mathbf{e}_i \cdot \mathbf{e}_j = \delta_{ij}$$

$$|\mathbf{E}_1|^2 = (J^2 - s^2) \left(\frac{n_0^2}{\lambda_0^2} - 4\kappa J^2 \right), \quad |\mathbf{E}_3|^2 = (J^2 - s^2) \left(1 - \frac{4s^2 \lambda_0^2 \kappa}{n_0^2} \right), \quad |\mathbf{E}_2|^2 = |\mathbf{E}_1|^2 |\mathbf{E}_3|^2$$

Decomposing \mathbf{r} over this frame, we introduce the polar coordinates:

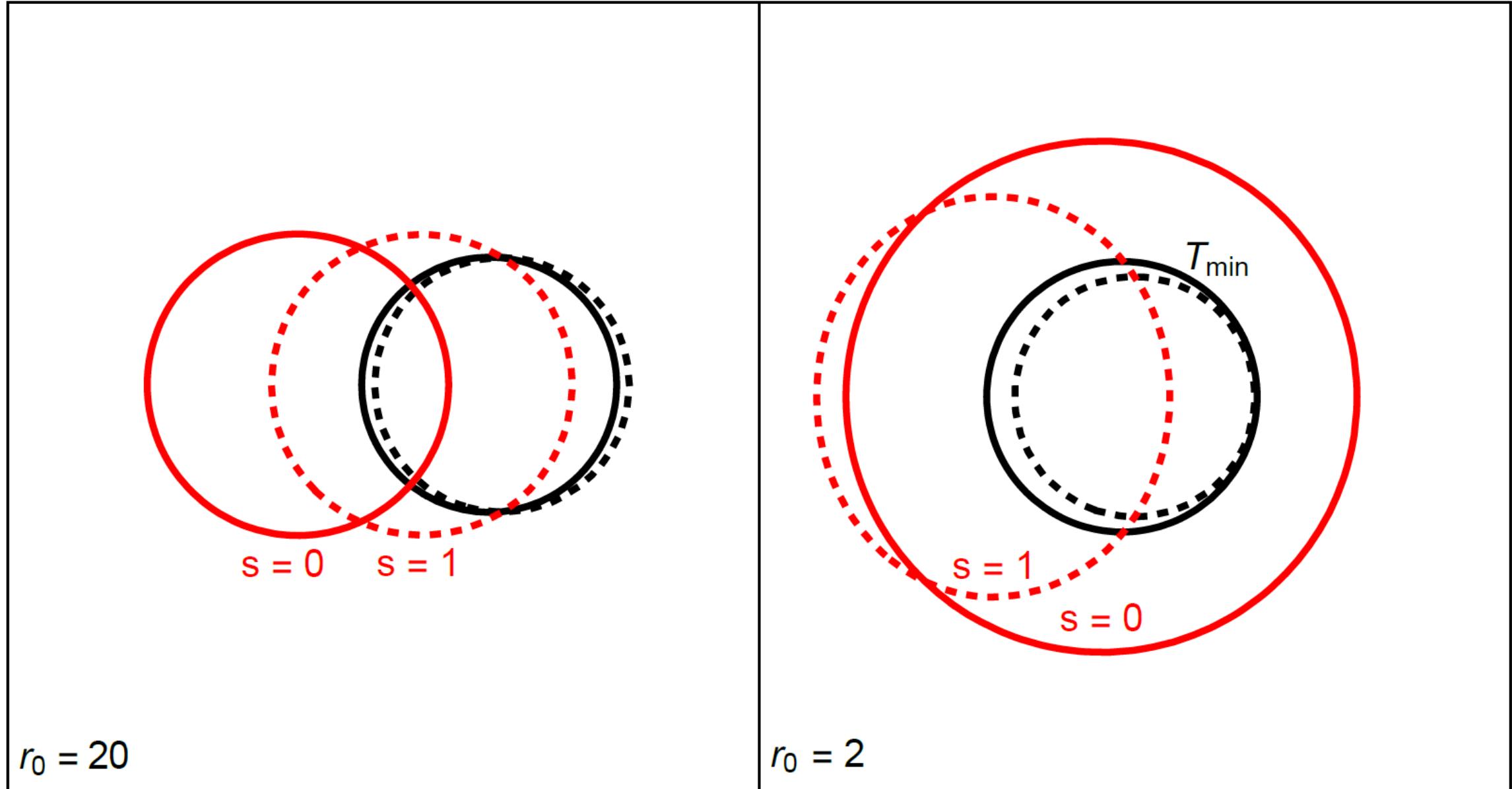
$\mathbf{r} = \underline{x}_1 \mathbf{e}_1 + \underline{x}_2 \mathbf{e}_2$, $\underline{x}_1 = r \cos \varphi$, $\underline{x}_2 = r \sin \varphi$ and we get the trajectory equation

$$1 - |\kappa| |a_s| r \cos \varphi = \frac{1 + \kappa r^2}{1 + \sqrt{1 - \frac{4\kappa s^2 \lambda_0^2}{n_0^2} (1 + \kappa r^2)}}$$

where

$$|a_s|^2 := R_s^2 - \frac{1}{\kappa}, \quad R_s^2 := \frac{n_0^2 - 4\kappa s^2 \lambda_0^2}{4\lambda_0^2 \kappa^2 (J^2 - s^2)}$$

Deformations of the trajectories



THANK YOU